

Jet Noise Predictions Based on Two Different Forms of Lilley's Equation

Part 2: Acoustic Predictions and Comparison With Data

Marvin E. Goldstein
Glenn Research Center, Cleveland, Ohio

Abbas Khavaran
QSS Group, Inc., Cleveland, Ohio

Ricardo E. Musafir
Federal University of Rio de Janeiro, Rio de Janeiro, Brazil

The NASA STI Program Office . . . in Profile

Since its founding, NASA has been dedicated to the advancement of aeronautics and space science. The NASA Scientific and Technical Information (STI) Program Office plays a key part in helping NASA maintain this important role.

The NASA STI Program Office is operated by Langley Research Center, the Lead Center for NASA's scientific and technical information. The NASA STI Program Office provides access to the NASA STI Database, the largest collection of aeronautical and space science STI in the world. The Program Office is also NASA's institutional mechanism for disseminating the results of its research and development activities. These results are published by NASA in the NASA STI Report Series, which includes the following report types:

- **TECHNICAL PUBLICATION.** Reports of completed research or a major significant phase of research that present the results of NASA programs and include extensive data or theoretical analysis. Includes compilations of significant scientific and technical data and information deemed to be of continuing reference value. NASA's counterpart of peer-reviewed formal professional papers but has less stringent limitations on manuscript length and extent of graphic presentations.
- **TECHNICAL MEMORANDUM.** Scientific and technical findings that are preliminary or of specialized interest, e.g., quick release reports, working papers, and bibliographies that contain minimal annotation. Does not contain extensive analysis.
- **CONTRACTOR REPORT.** Scientific and technical findings by NASA-sponsored contractors and grantees.

- **CONFERENCE PUBLICATION.** Collected papers from scientific and technical conferences, symposia, seminars, or other meetings sponsored or cosponsored by NASA.
- **SPECIAL PUBLICATION.** Scientific, technical, or historical information from NASA programs, projects, and missions, often concerned with subjects having substantial public interest.
- **TECHNICAL TRANSLATION.** English-language translations of foreign scientific and technical material pertinent to NASA's mission.

Specialized services that complement the STI Program Office's diverse offerings include creating custom thesauri, building customized databases, organizing and publishing research results . . . even providing videos.

For more information about the NASA STI Program Office, see the following:

- Access the NASA STI Program Home Page at <http://www.sti.nasa.gov>
- E-mail your question via the Internet to help@sti.nasa.gov
- Fax your question to the NASA Access Help Desk at 301-621-0134
- Telephone the NASA Access Help Desk at 301-621-0390
- Write to:
NASA Access Help Desk
NASA Center for Aerospace Information
7121 Standard Drive
Hanover, MD 21076



Jet Noise Predictions Based on Two Different Forms of Lilley's Equation

Part 2: Acoustic Predictions and Comparison With Data

Marvin E. Goldstein
Glenn Research Center, Cleveland, Ohio

Abbas Khavaran
QSS Group, Inc., Cleveland, Ohio

Ricardo E. Musafir
Federal University of Rio de Janeiro, Rio de Janeiro, Brazil

Prepared for the
34th International Congress and Exposition on Noise Control Engineering
cosponsored by CAPES, Isover, Ecophon, Eurocoustic, CertainTeed, 3M,
01dB-Metravib, and Embraer
Rio de Janeiro, Brazil, August 7–10, 2005

National Aeronautics and
Space Administration

Glenn Research Center

Acknowledgments

The authors would like to thank Nicholas Georgiadis, Nozzle Branch, NASA Glenn Research Center, for providing the RANS-Wind solution to several jets discussed in this paper.

This report is a formal draft or working paper, intended to solicit comments and ideas from a technical peer group.

This report contains preliminary findings, subject to revision as analysis proceeds.

This report is a preprint of a paper intended for presentation at a conference. Because of changes that may be made before formal publication, this preprint is made available with the understanding that it will not be cited or reproduced without the permission of the author.

Available from

NASA Center for Aerospace Information
7121 Standard Drive
Hanover, MD 21076

National Technical Information Service
5285 Port Royal Road
Springfield, VA 22100

Available electronically at <http://gltrs.grc.nasa.gov>

Jet Noise Predictions Based on Two Different Forms of Lilley's Equation

Part 2: Acoustic Predictions and Comparison with Data

Marvin E. Goldstein
National Aeronautics and Space Administration
Glenn Research Center
Cleveland, Ohio 44135

Abbas Khavaran
QSS Group, Inc.
Cleveland, Ohio 44135

Ricardo E. Musafir
Universidade Federal do Rio de Janeiro
Rio de Janeiro 21945-970, Brazil

Abstract

The far field acoustic spectra at 90° to the downstream axis of some typical high speed jets are calculated from two different forms of Lilley's equation combined with some recent measurements of the relevant turbulent source function. These measurements, which were limited to a single point in a low Mach number flow, were extended to other conditions with the aid of a highly developed RANS calculation. The results are compared with experimental data over a range of Mach numbers. Both forms of the analogy lead to predictions that are in fair agreement with the experimental data at subsonic Mach numbers. The agreement is not quite as good at supersonic speeds, but the data appears to be slightly contaminated by shock- associated noise in this case.

I. Introduction

The results of Part 1 can not be used to predict the radiated sound without inputting more specific information about the turbulence structure. In this part we accomplish this objective by using some recent measurements (ref. 1) of the two point fourth order stream-wise velocity correlation spectra that were carried out by Harper-Bourne (ref. 1) at a single point on the centerline of the mixing layer in a low Mach number jet.

II. The Harper-Bourne Spectrum

Harper-Bourne's results would most closely correspond to

$$H_o(y, \eta, \omega) \equiv \frac{1}{\pi} \int_{-\infty}^{\infty} e^{-i\omega\tau} R_{11}^2(y, \eta, \tau) d\tau \quad (1)$$

with the quasi-normal approximation that is being used in the present analysis.

He divided this quantity into the three components (see his eq. (2.5) and (2.7) on p. 2)

$$H_o = H_o(\mathbf{y}, \mathbf{0}, \omega) R\left(\mathbf{y}, \frac{\eta_{\parallel}}{l_{\parallel}}, \frac{\eta_{\perp}}{l_{\perp}}, \omega\right) e^{i\omega\tau_p} \quad (2)$$

where l_{\parallel} , l_{\perp} are the spectral stream-wise and transverse length scales (not to be confused with the time domain length scales L_{\parallel} and L_{\perp} introduced above) and

$$\tau_p \approx \frac{\eta_{\parallel}}{U_c} \quad (3)$$

No assumption is made about the decomposition of the correlations into products of their space and time components with this approach.

The first factor can be evaluated from his measurements of $R_{1111}(\mathbf{y}, \mathbf{0}, \tau)$, which can be reasonably well represented by the simple exponential $e^{-\lambda|\tau|}$. Taking its Fourier Transform shows that (ref. 20)

$$H_o(\mathbf{y}, \mathbf{0}, \omega) = \frac{\widetilde{\lambda \rho u_1^4}}{\pi(\lambda^2 + \omega^2)} \quad (4)$$

Harper-Bourne was able to obtain a reasonable fit to his data with the non-separable form

$$R = e^{-\sqrt{\bar{\eta}_{\parallel}^2 + \bar{\eta}_{\perp}^4}}, \quad (5a)$$

$\bar{\eta}_{\parallel} \equiv \eta_{\parallel}/l_{\parallel}$ and $\bar{\eta}_{\perp} \equiv \eta_{\perp}/l_{\perp}$ by allowing the stream-wise and transverse length scales l_{\parallel} and l_{\perp} to be frequency dependant. A better fit might be

$$R = e^{-\sqrt{\bar{\eta}_{\parallel}^2 + \bar{\eta}_{\perp}^4}} \left(1 - \frac{b\bar{\eta}_{\parallel}^2}{\sqrt{\bar{\eta}_{\parallel}^2 + \bar{\eta}_{\perp}^4}} \right) \quad (5b)$$

where

$$b = \frac{1 + \beta^2}{1 + 4\beta^2} \quad (6)$$

with $\beta \equiv \pi U_J / U_c$.

Inserting these into equation (2) and using the result in equations (43) and (44) of Part 1 shows that

$$I_{\omega}(\mathbf{x}|\mathbf{y}) = C_0^2 \frac{\widetilde{\lambda \rho u_1^4} \left[\omega^2 + (\kappa |\nabla U|)^2 \right] \pi U_c^3 \bar{l}_{\parallel} \bar{l}_{\perp}^2}{x^2 c_{\infty}^4 (\lambda^2 + \omega^2) \omega \left[1 + (2\pi \bar{l}_{\parallel})^2 \right]^{3/2}} \left[1 - b \frac{1 + 4(2\pi \bar{l}_{\parallel})^2}{1 + (2\pi \bar{l}_{\parallel})^2} \right], \quad (7)$$

Where the scaled length scales \bar{l}_{\parallel} and \bar{l}_{\perp} are defined by

$$\bar{l}_1 \equiv \frac{\omega l_1}{2\pi U_c} \quad (8a)$$

and

$$\bar{l}_\perp \equiv \frac{\omega l_\perp}{2\pi U_c} \quad (8b)$$

Harper-Bourne's figure 13 shows that while l_1 and l_\perp are constant at relatively low frequencies it is the scaled length scales \bar{l}_1 and \bar{l}_\perp that becoming constant as $\omega \rightarrow \infty$. The data is reasonably well represented by the functions

$$\frac{\omega l_1}{2\pi U_J} \approx \frac{1}{2} (1 - e^{-2St}) \quad (9a)$$

and

$$\frac{\omega l_\perp}{2\pi U_J} \approx 0.15 (1 - e^{-0.5St}) \quad (9b)$$

where

$$St \equiv \omega D / 2\pi U_J \quad (10)$$

As indicated in Part 1, this result was derived only for the first formulation, but it turns out that it will apply to the second as well if κ is set to zero and a slightly different formula is used for C_0^2 . The principle difference between these results is therefore due to the factor $\left[\omega^2 + (\kappa |\nabla U|)^2 \right]$, which does not significantly effect the high frequency behavior of the solution but causes I_ω to exhibit the dipole-like behavior

$$I_\omega \sim \omega^2 \quad \text{as } \omega \rightarrow 0 \quad (11)$$

in the first formulation and the quadrupole-like behavior

$$I_\omega \sim \omega^4 \quad \text{as } \omega \rightarrow 0 \quad (12)$$

in the second.

III. Extension of the Harper-Bourne Data

Unfortunately, all of Harper-Bourne's measurements were taken at a single point in a very low Mach number jet, while acoustic predictions require information about the turbulence over the entire noise producing region of the jet. We, therefore attempt to extend his data by using some modeling assumptions along with the Glenn Wind code, which is a RANS code with a standard $k - \epsilon$ turbulence model. To this

end, we first assume that the time scale λ^{-1} that appears in equation (53) of Part 1 is proportional to the $k - \varepsilon$ time scale k/ε , i.e., we put

$$\lambda^{-1} \approx C^\tau k/\varepsilon \quad (13)$$

where C^τ is an adjustable constant. In order to extend equations (9a) and (9b), we assume that the time and velocity scales D/U_J and U_J are proportional to the $k - \varepsilon$ time and length scales k/ε and $k^{1/2}$ respectively to obtain

$$\frac{\omega l_1}{2\pi U_J} \approx C^l k^{1/2} \left[1 - e^{-C^S \left(\frac{\omega k}{2\pi \varepsilon} \right)} \right] \quad (14a)$$

and

$$\frac{\omega l_1}{2\pi U_J} \approx 0.3 C^l k^{1/2} \left[1 - e^{-0.25 C^S \left(\frac{\omega k}{2\pi \varepsilon} \right)} \right] \quad (14b)$$

where the constants C^l and C^S are determined by requiring that equations (9a) and (9b) be in agreement with Harper-Bourne's measurements at the Harper-Bourne measuring point and Mach number when k and ε are calculated from the Wind code. A reasonably good approximation is obtained by putting $C^l \approx 3.3$ and $C^S \approx 0.40$. To be consistent with these extensions it is necessary to put

$$\beta = \frac{2\pi C^l k^{1/2}}{U_c} \quad (15)$$

in equation (6).

In the first formulation the constant C_o is related to the ratio Γr (defined in equation (42) of Part 1 by equation (45) of Part 1), which for $\gamma = 1.4$ becomes

$$C_o^2 \approx \frac{0.43}{3} (\Gamma r)^2 + 0.01 \quad (16a)$$

and in the second by

$$C_o^2 = \frac{1}{4} (\Gamma r)^2 \quad (16b)$$

Unfortunately, Harper-Bourne only measured the stream-wise and not the transverse velocity correlations so that Γ is essentially unknown. We therefore treat C^τ and C_o as adjustable constants, whose determination is described in the next section. It is necessary to know the square root in equation (46) of Part 1 in order to determine κ , but again, Harper-Bourne does not provide enough data to determine this quantity. We estimate its value to be less than one, however.

IV. Comparison with acoustic Measurements and Discussion

RANS solutions for Mach 0.50, 0.90, and 1.5 cold jets were obtained from the WIND code with upstream nozzle conditions specified in terms of plenum temperature ratio T_r and the pressure ratio. The predicted turbulent kinetic energy distributions for the three jets are shown in figure 1.

The far-field spectra at 90° to the jet axis were calculated for these jets on the arc $R/D = 100$ by summing the point result (7) over the noise producing region of the jet. Figures 2 through 4 show the comparison between these results and the subsonic SHJAR data recently acquired at NASA Glenn Research Center and correctly expanded supersonic data obtained at Langley Research Center. Atmospheric attenuation was removed from all measurements in order to make lossless comparisons with the predictions. The agreement is better in the subsonic case, but it is likely that the supersonic data contains a small amount of shock associated noise that is not accounted for by the theory.

The adjustable constants C^T and C_o were determined by obtaining the best fit with the Mach 0.5 data. The resulting value of C^T turns out to be 0.10. Figure 3 shows that there is almost no dependence on the parameter κ when its value is in the estimated range $0 < \kappa < 1$, which means that these data comparisons were not discriminating enough to distinguish between these two formulations. The hope is that similar comparisons for hot jets or jets with more complex flow fields will provide the required selectivity.

V. Concluding Remarks

The research was initially motivated by the desire to distinguish between the two forms of the acoustic analogy described above. Unfortunately the results turned out to be inconclusive—with both forms of the analogy yielding reasonable agreement with the data. Our hope is that similar comparisons for hot jets or jets with more complex flow fields will provide the required selectivity. But until this is done, our recommendation would be to base the jet noise predictions on the second formulation, since it leads to much simpler formulas at angles other than 90° .

References

1. Harper-Bourne, M., “Jet Noise Turbulence Measurements,” AIAA Paper 2003–3214, 2003.
2. Pope, S.B., “Turbulent Flows,” Cambridge University Press, 2000, pp. 70 and 144.
3. Khavaran, A., Bridges, J., and Freund, J.B., “A Parametric Study of Fine-Scale Turbulence Mixing Noise,” NASA/TM—2002-211696, 2002
4. Norum, T., NASA Langley Jet Noise Lab Generic Nozzle Narrow Band Spectra, vol. 1, 1994 (unpublished, available electronically).

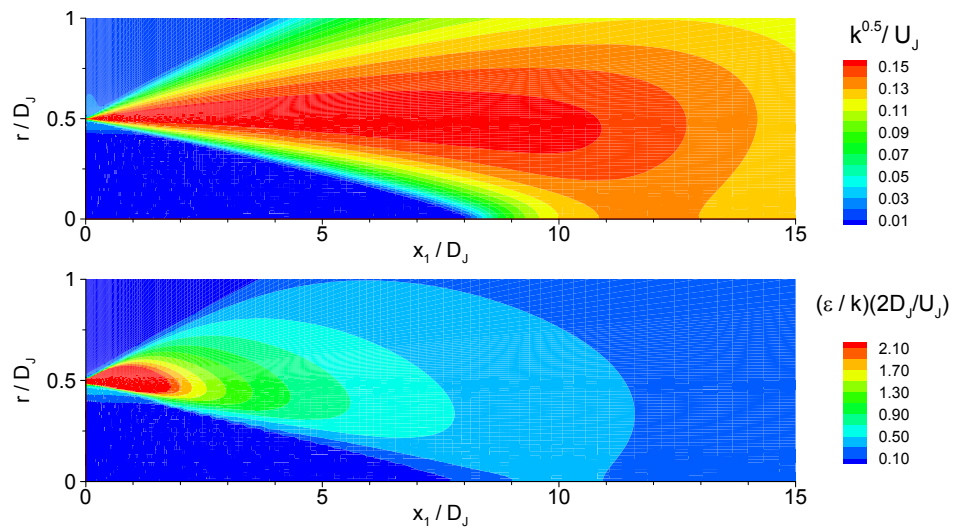


Figure 1(a).—As figure 1 but Predicted turbulent kinetic energy (top), and frequency scale (bottom) for a 2 in. diameter cold jet at Mach 0.50 $r \equiv |y_\perp|$.

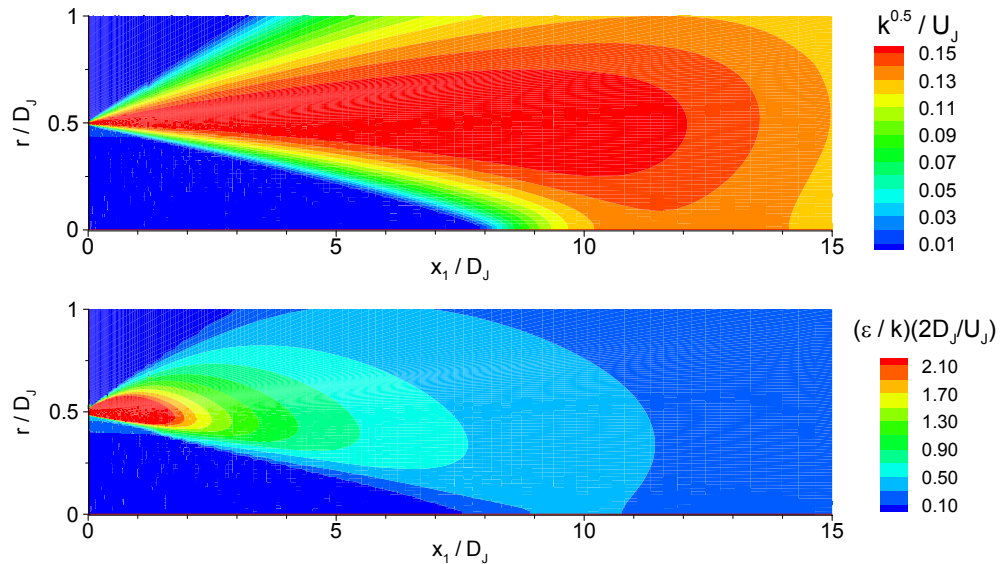


Figure 1(b).—As figure 1 but predicted turbulent kinetic energy (top), and frequency scale (bottom) for a 2 in. diameter cold jet at Mach 0.90 $r \equiv |y_\perp|$.

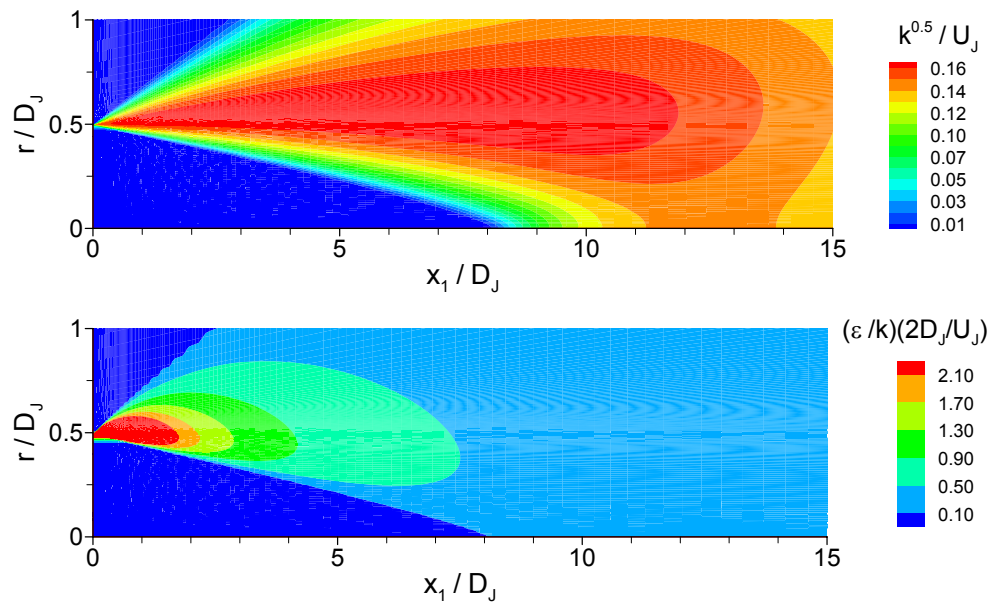


Figure 1(c).—As figure 1 but predicted turbulent kinetic energy (top), and frequency scale (bottom) for a Mach 1.50 convergent-divergent nozzle with 1.68 in. exit diameter $r \equiv |y_{\perp}|$.

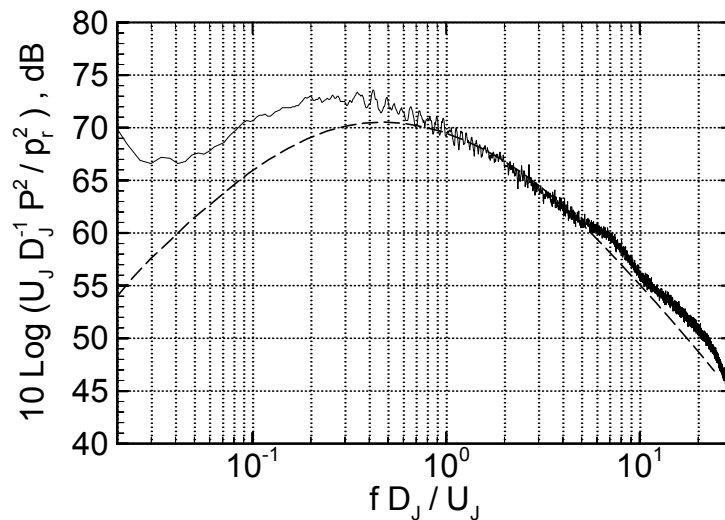


Figure 2.—Spectrum at 90° and at $R/D = 100$ for a Mach 0.50 cold jet. Prediction (dashed line); data (solid line).

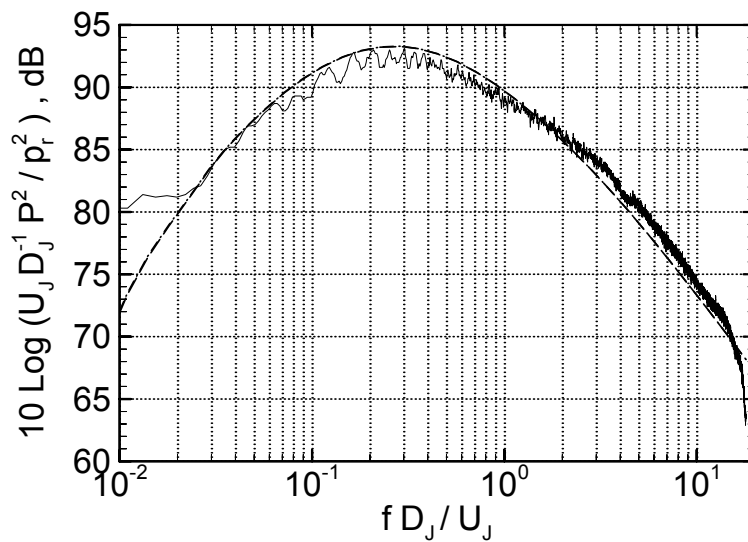


Figure 3.—As figure 1 but for a Mach 0.90 cold jet. Prediction with $\kappa = 0.0$ (dashed line); $\kappa = 0.90$ (dash-dot); data (solid line).

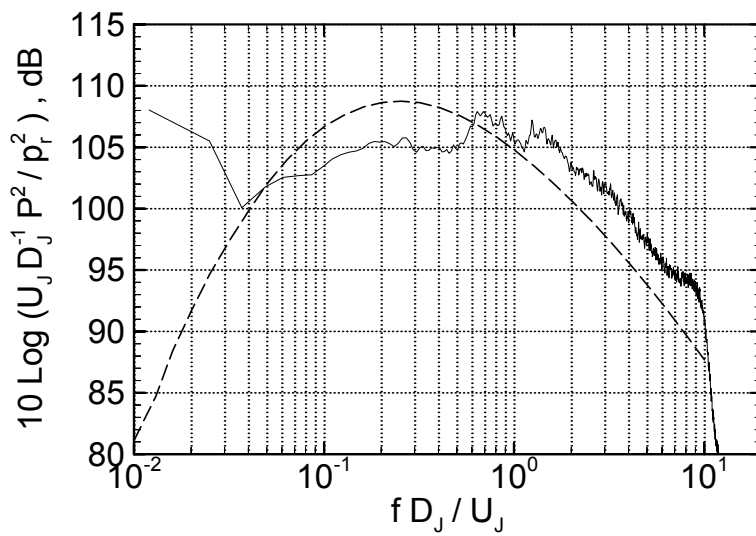


Figure 4.—As figure 2 but for Mach 1.5 cold jet.

REPORT DOCUMENTATION PAGE			Form Approved OMB No. 0704-0188	
Public reporting burden for this collection of information is estimated to average 1 hour per response, including the time for reviewing instructions, searching existing data sources, gathering and maintaining the data needed, and completing and reviewing the collection of information. Send comments regarding this burden estimate or any other aspect of this collection of information, including suggestions for reducing this burden, to Washington Headquarters Services, Directorate for Information Operations and Reports, 1215 Jefferson Davis Highway, Suite 1204, Arlington, VA 22202-4302, and to the Office of Management and Budget, Paperwork Reduction Project (0704-0188), Washington, DC 20503.				
1. AGENCY USE ONLY (Leave blank)		2. REPORT DATE September 2005		3. REPORT TYPE AND DATES COVERED Technical Memorandum
4. TITLE AND SUBTITLE Jet Noise Predictions Based on Two Different Forms of Lilley's Equation Part 2: Acoustic Predictions and Comparison With Data			5. FUNDING NUMBERS Cost Center 2250000002	
6. AUTHOR(S) Marvin E. Goldstein, Abbas Khavaran, and Ricardo E. Musafir				
7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES) National Aeronautics and Space Administration John H. Glenn Research Center at Lewis Field Cleveland, Ohio 44135-3191			8. PERFORMING ORGANIZATION REPORT NUMBER E-15199-2	
9. SPONSORING/MONITORING AGENCY NAME(S) AND ADDRESS(ES) National Aeronautics and Space Administration Washington, DC 20546-0001			10. SPONSORING/MONITORING AGENCY REPORT NUMBER NASA TM-2005-213829-PART2	
11. SUPPLEMENTARY NOTES Prepared for the 34th International Congress and Exposition on Noise Control Engineering cosponsored by CAPES, Isover, Ecophon, Eurocoustic, CertainTeed, 3M, 01dB-Metravib, and Embraer, Rio de Janeiro, Brazil, August 7-10, 2005. Marvin E. Goldstein, NASA Glenn Research Center; Abbas Khavaran, QSS Group, Inc., 21000 Brookpark Road, Cleveland, Ohio 44135; and Ricardo E. Musafir, Federal University of Rio de Janeiro, Rio de Janeiro, Brazil. Responsible person, Marvin E. Goldstein, organization code RO, 216-433-5825.				
12a. DISTRIBUTION/AVAILABILITY STATEMENT Unclassified - Unlimited Subject Category: 71 Available electronically at http://gltrs.grc.nasa.gov This publication is available from the NASA Center for AeroSpace Information, 301-621-0390.			12b. DISTRIBUTION CODE	
13. ABSTRACT (Maximum 200 words) The formulations of the acoustic analogy developed in Part 1 are combined with some recent measurements of the appropriate turbulent source function in order to calculate the far field acoustic spectra at 90° to the downstream axis of some typical high speed jets. The source measurements, which were limited to a single point in a low Mach number flow, were extended to other conditions with the aid of a RANS code with a k -e turbulence model. The results are compared with experimental data over a range of Mach numbers. Both forms of the analogy lead to predictions that are in fair agreement is not quite as good at supersonic speeds, but the data appears to be slightly contaminated by shock-associated noise in this case.				
14. SUBJECT TERMS Aeroacoustics			15. NUMBER OF PAGES 14	
			16. PRICE CODE	
17. SECURITY CLASSIFICATION OF REPORT Unclassified	18. SECURITY CLASSIFICATION OF THIS PAGE Unclassified	19. SECURITY CLASSIFICATION OF ABSTRACT Unclassified	20. LIMITATION OF ABSTRACT	

